

Radiation

M. Burba | Numerical Model Training | July 2025

Material: S. Schäfer, R. Hogan, M. Ahlgrimm

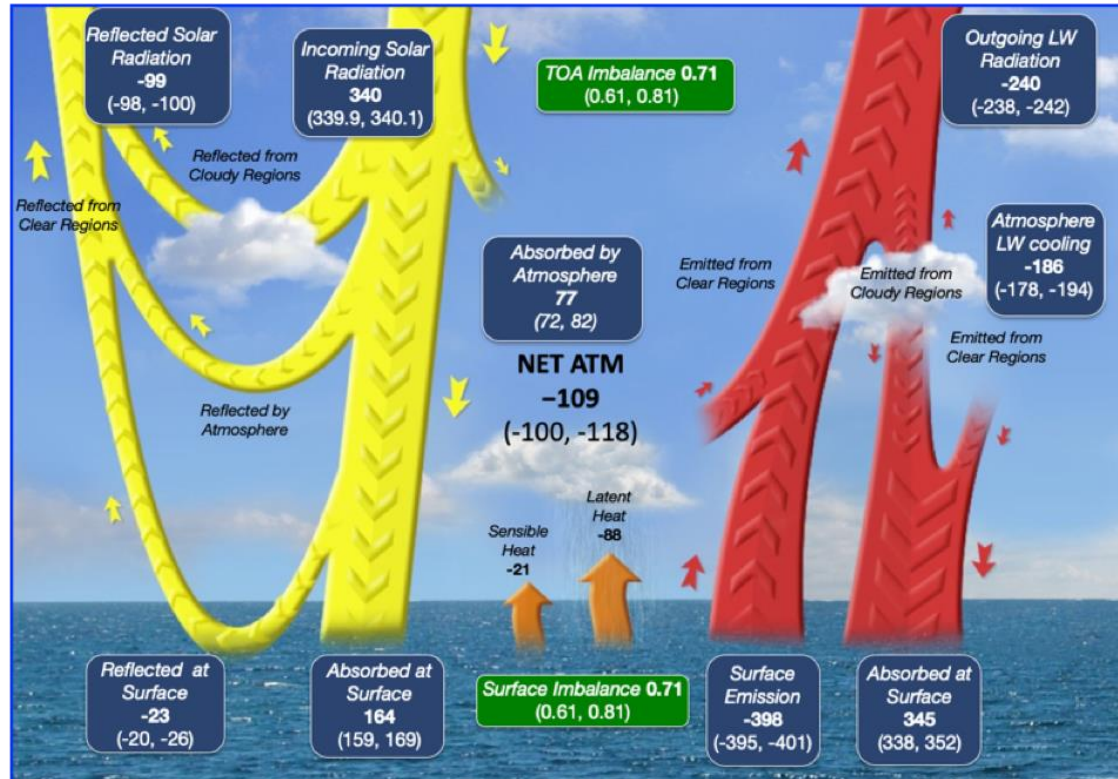


- Physics of radiation in the atmosphere
 - Motivation: why do we care?
 - Bulk radiation treatment
 - Spectra and atmospheric gases
 - Scattering by atmospheric particles
 - Aerosol and surface properties
- Radiation modelling (in ICON)
 - Specifically: ecRAD
- How to run radiation
 - Namelist settings
 - Uncertainties

Introduction

Radiation – why do we care?



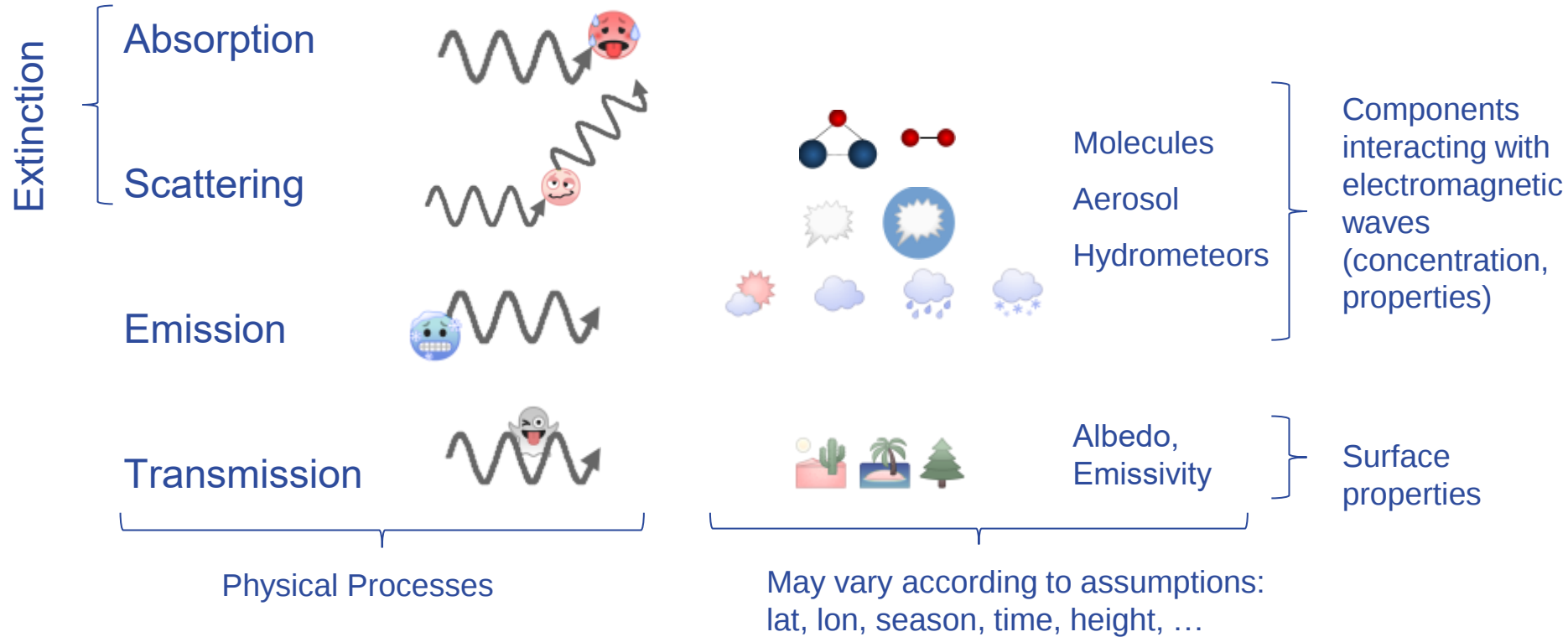


“Average view”

Balance between SW heating and LW cooling ultimately determines average surface temperature

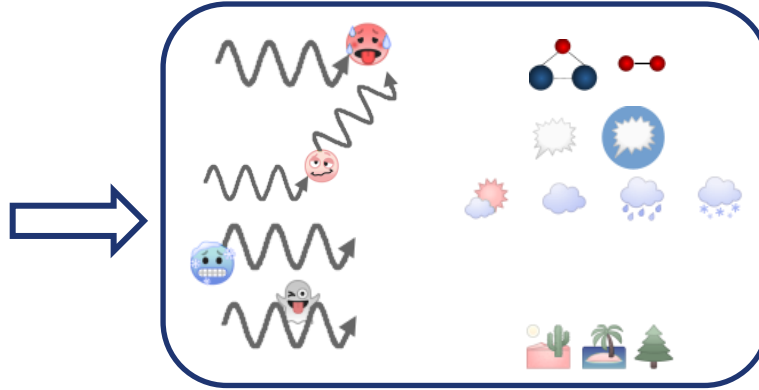
Source:
<https://ceres.larc.nasa.gov/images/EnergyBudget.png>

- Differential heating at poles and equator drives large-scale dynamics
 - Pole: radiative cooling, heated by circulation
 - Equator: effective heating by the sun, cooled by circulation
- Heating (cooling) in low layers (high layers) drives convection



Inputs for ecRad:

- Atmospheric fields:
Temperature, Pressure,
water vapor,
hydrometeors, cloud
cover
- Aerosol, surface albedo,
surface emissivity, T_{skin} ,
land fraction, glacier
fraction, effective radius
- Sun zenith angle
- GHG concentrations



ecRad

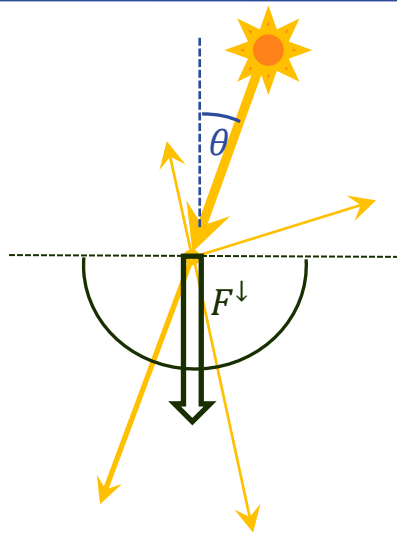
Outputs of ecRad:

- Fluxes, SW and LW
- VIS, NIR, PAR

Physics of radiation

Bulk Radiation Treatment

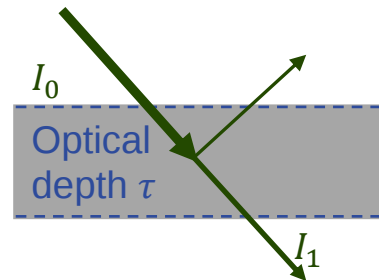




- **Radiance R** : power per unit area and solid angle
- **Flux F** : power per area integrated over angles, e.g. total up- or downward flux
- $F^\downarrow(\lambda) = \int_{\Omega} R_\lambda d\Omega$, λ wavelength

Bulk approximation:
Upwelling/downwelling flux only!

- **Extinction coefficient β** : local extinction by absorption and scattering
- **Optical depth $\tau = \beta s$** measures extinction along a path s
- **Transmission $T = 1 - \exp(-\tau)$** amount of incoming radiation which is not scattered or absorbed by a layer



$$\frac{dF_{diff}^{\uparrow}}{d\tau} = F_{diff}^{\uparrow} - F_{diff}^{\downarrow} - \text{sources}$$

$$\frac{dF_{diff}^{\downarrow}}{d\tau} = F_{diff}^{\downarrow} - F_{diff}^{\uparrow} + \text{sources}$$

$$\frac{dF_{direct}^{\downarrow}}{d\tau} = -\text{sources}$$

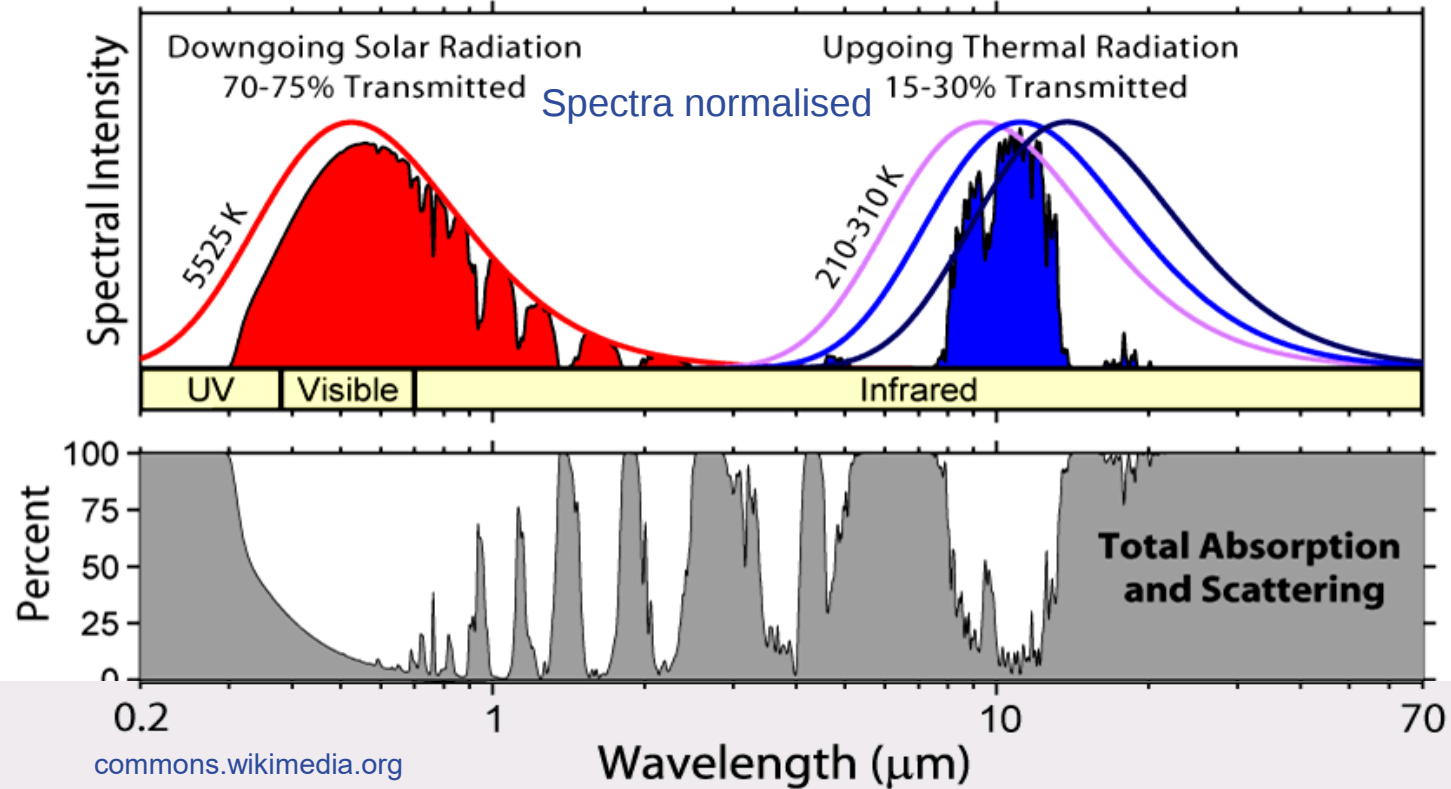
Note: ignored factors...

Physics of Radiation

Radiation spectra and atmospheric gases



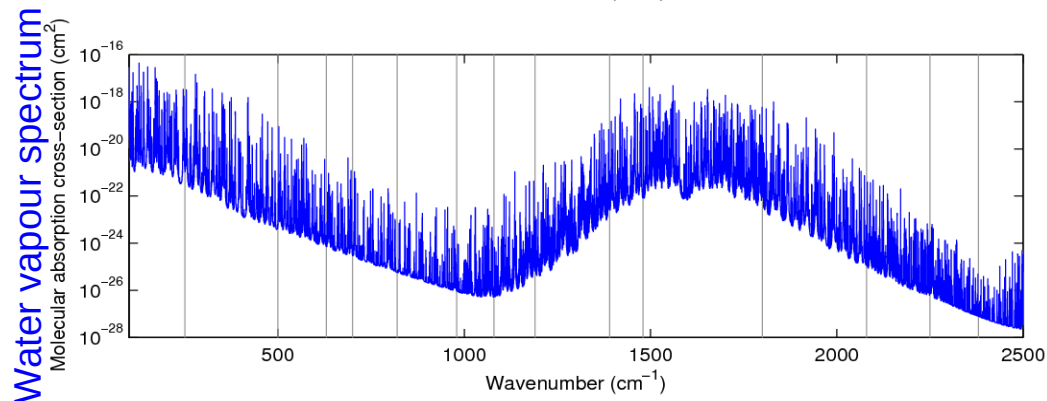
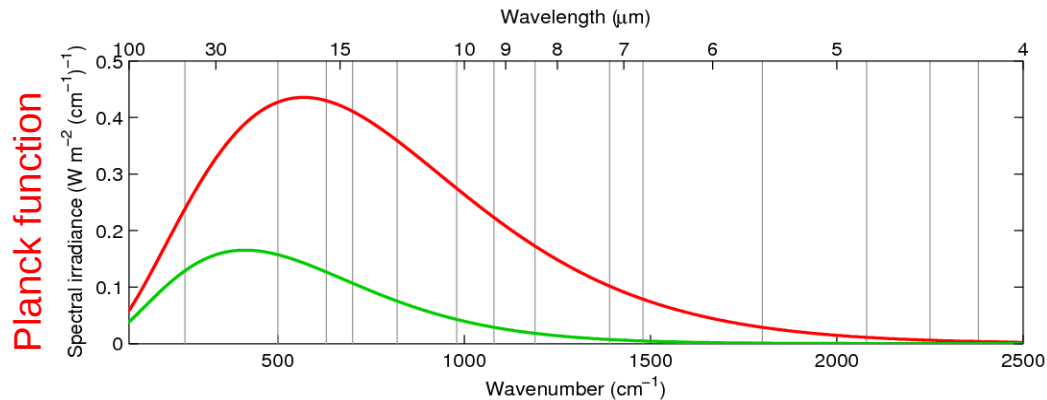
ICON Incoming SW and outgoing LW radiation spectra



Gas	Parts by volume (per million)	Interaction
Nitrogen (N ₂)	780,840 ppm (78.084%)	sw (scat)
Oxygen (O ₂)	209,460 ppmv (20.946%)	sw (scat, abs)
Water vapour (H ₂ O)	~0.40% total, surface ~1%-4%	lw, sw (abs)
Argon (Ar)	9,340 ppmv (0.9340%)	
Carbon dioxide (CO ₂)	390 ppmv (0.039%) rising	lw, sw (abs)
Neon (Ne)	18.18 ppmv (0.001818%)	
Helium (He)	5.24 ppmv (0.000524%)	
Methane (CH ₄)	1.79 ppmv (0.000179%) rising	lw
Krypton (Kr)	1.14 ppmv (0.000114%)	
Hydrogen (H ₂)	0.55 ppmv (0.000055%)	
Nitrous oxide (N ₂ O)	0.319 ppmv (0.00003%) rising	lw
Carbon monoxide (CO)	0.1 ppmv (0.00001%)	
Xenon (Xe)	0.09 ppmv (0.000009%)	
Ozone (O ₃)	0.0 to 0.07 ppmv (0.000007%)	lw, sw (abs)

- Gas-radiation interactions: shortwave (sw) absorption (abs) and scattering (scat), longwave (lw) absorption (greenhouse gases)
- Not all gases radiatively active, small concentrations can have strong effect → GHG

- “Rapid Radiative Transfer Model” gas-optics
- Divide spectrum into bands where Planck function is similar
- ICON uses RRTMG (Mlawer et al. 1997, Iacono et al. 2008): 14 bands in shortwave, 16 in longwave



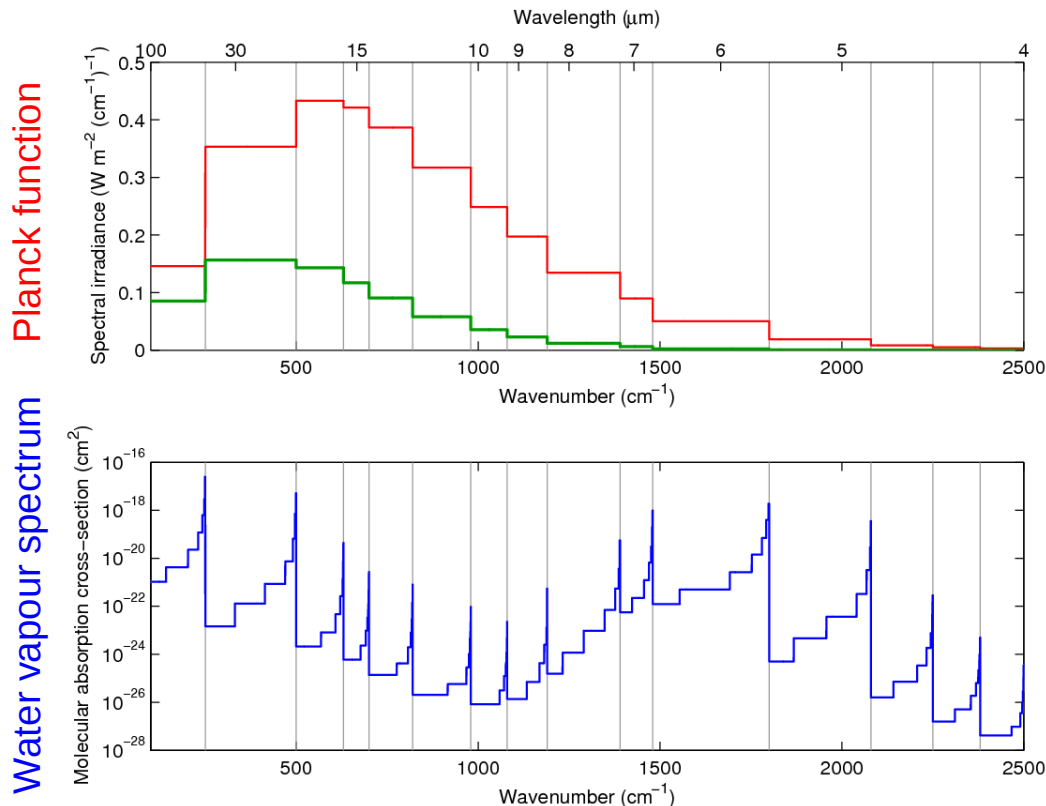
■ In each **band**:

- Approximate Planck function
- Re-order by gas absorption, approximate in 6-21 **g-points** (Lacis, Oinas 1991)

■ RRTMG: ~200 g-points

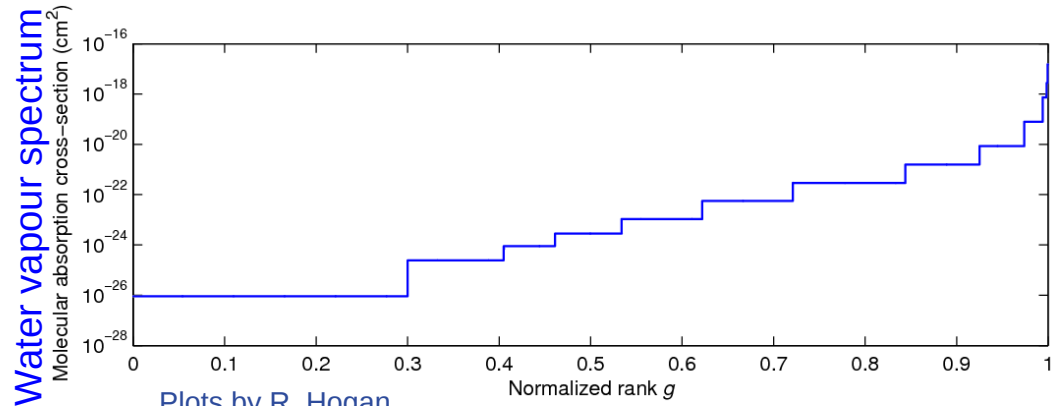
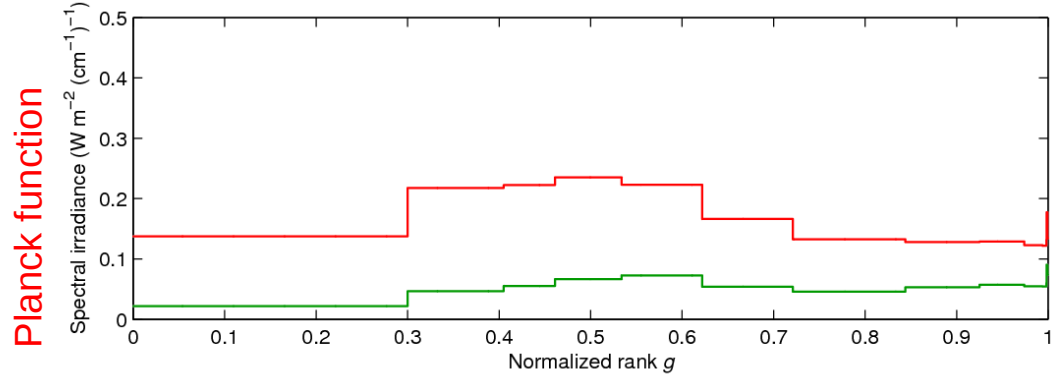
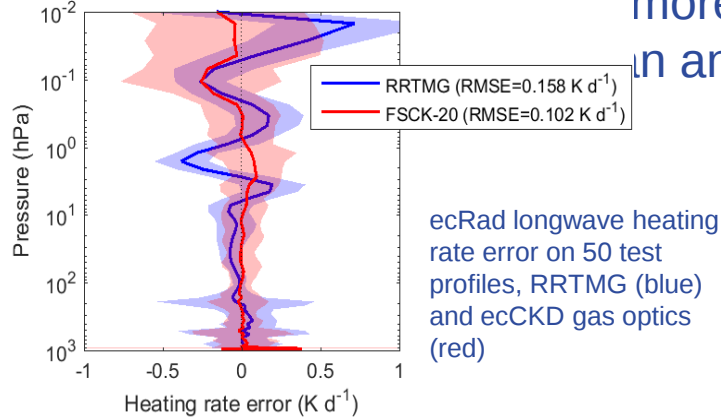
■ Gas on g-points, particles/surface on bands

■ Drawback: many repeated values in spectrum



- Re-order whole spectrum, average Planck emission for wavelengths in each g-point, optics of cloud etc. can be per g-point or coarser band

With 64 g-points: cheaper, more accurate and



Plots by R. Hogan

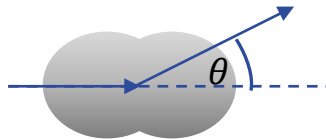
- Interaction of gases with radiation depends on the wavelength
- Full wavelength spectra must be simplified for computational efficiency
- Trade-off between accuracy and number of bands
- ecCKD: needs fewer bands to be more accurate than RRTMG
- DWD NEC vector machine: no immediate gain on computational time by reducing bands from ~200 down to 64
 - Efficiency depends on hardware architecture!

Physics of Radiation

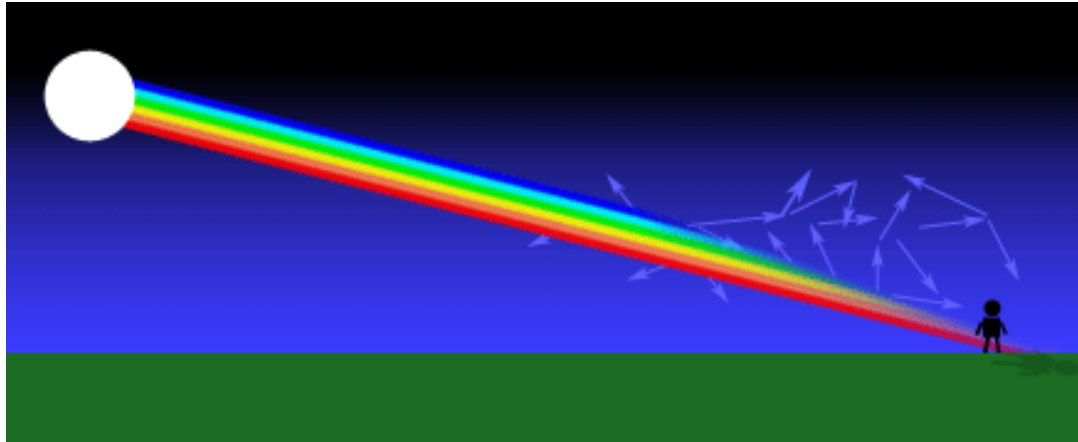
Scattering by atmospheric particles



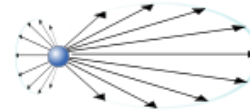
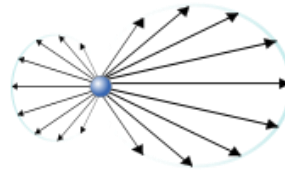
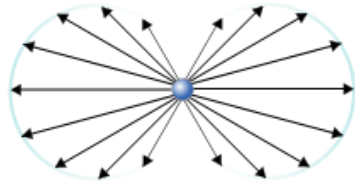
- Scattering intensity is a function of scattering angle θ and size parameter x (ratio of particle radius r and wavelength λ)
- $r \gg \lambda$: Geometric optics
- $r \ll \lambda$: Rayleigh scattering
 - Elastic scattering by spheres
 - Scattering intensity



- $p(\theta) = \frac{3}{4}(1 + (\cos \theta)^2)$
- Strong λ dependency: λ^{-4}
- Examples: blue sky during the day, red sunrise & sunset



Source:
<https://commons.wikimedia.org/wiki/File:RayleighScattering.gif>



Scattering intensity (θ , x)

Source:

https://upload.wikimedia.org/wikipedia/commons/thumb/6/67/Mie_scattering.svg/721px-Mie_scattering.svg.png?uselang=fr

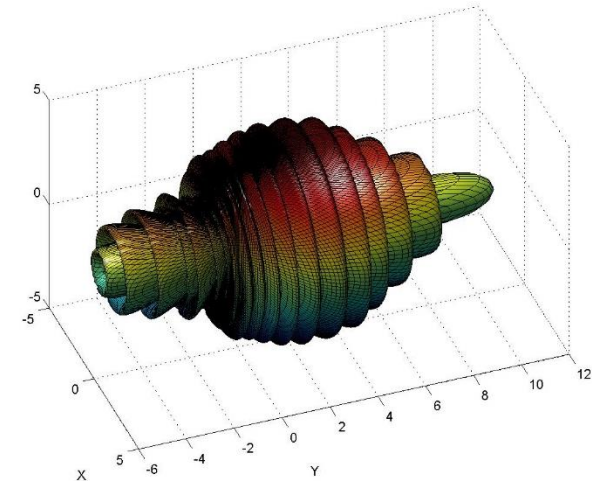
- $r \sim \lambda$, spherical particles : **Mie scattering**

- complex function of scattering angle
- Approximated by numerical algorithms
- Strong forward peak: treated together with non-scattered direct radiation in model

- Model: Simplify by bulk **asymmetry parameter g**

$g > 0$: more forward scattering

$g < 0$: more backward scattering



Source:

[3D_Miestreuung_an_2um_Kugel.jpg \(1201x900\)](#)

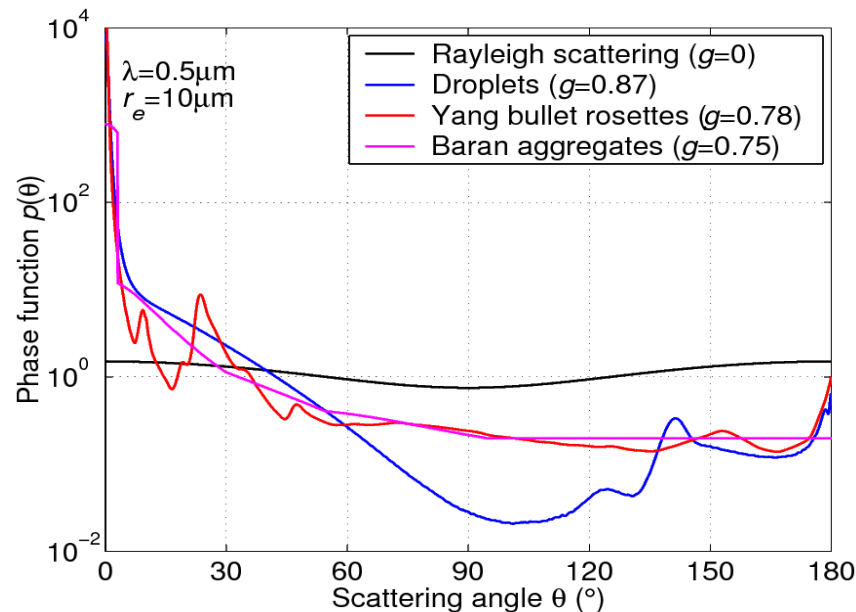
- Particle shape adds more complexity
- Ice particle shapes are complex
→ use shape assumptions; default in ICON: hexagonal columns (Fu 1996, 1998)



- Simplify by bulk **asymmetry parameter**

$$g = \frac{1}{4\pi} \int_0^\pi \int_0^{2\pi} p(\theta) \cos(\theta) d\phi d\theta ,$$

- $g > 0$: more forward scattering
- $g < 0$: more backward scattering



Plot by R. Hogan

- Clouds contain a mixture of particles of different sizes and shapes
- Distributions scatter like particles with one **effective radius**
 - r_{eff} = average radius weighted by number, area and scattering efficiency of each particle size
- Model: parametrise r_{eff}

- Size- and shape-dependent
- Heavily simplified in the model – large uncertainty
 - Complex shapes (ice) more uncertain than simpler shapes (liquid drops)
 - Assumptions made for radiation not necessarily the same as those for microphysics

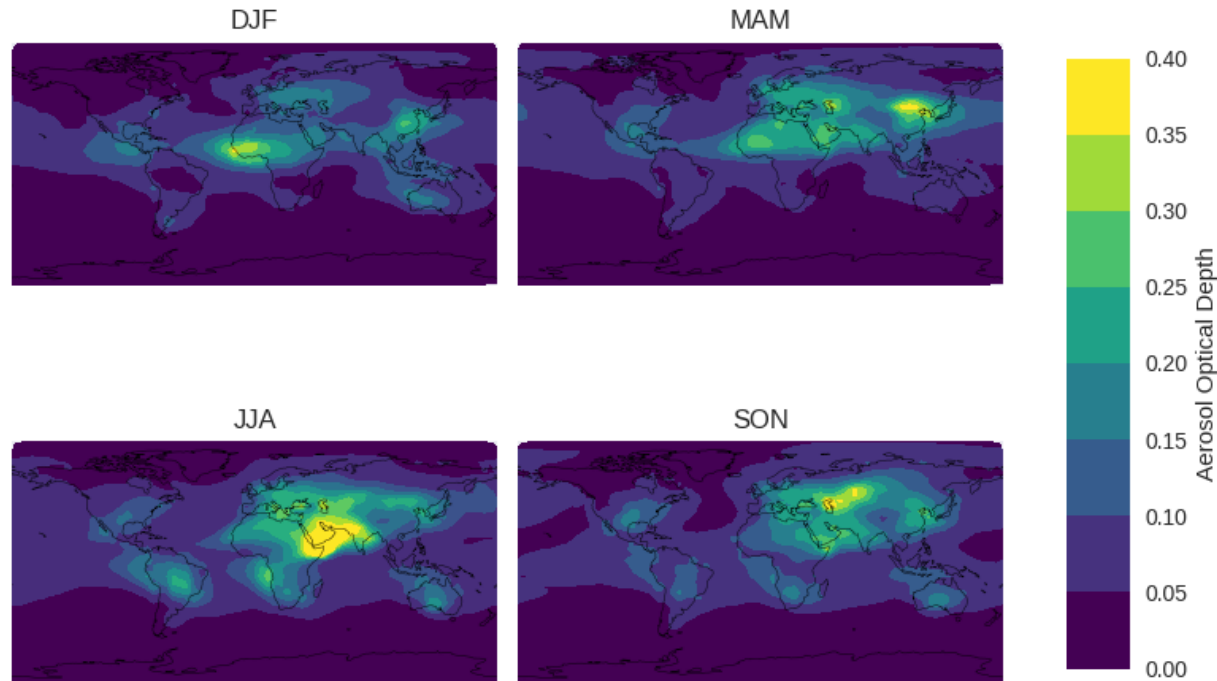
Physics of Radiation

Aerosol and surface properties



- Use aerosol climatology of optical properties e.g. Tegen et al 1997
- Tegen is operational NWP default
 - `extpar` file
 - `irad_aero = 6`

Tegen Aerosol (from extpar)



- Currently dynamic area of development!
- ICON-ART aerosol (*irad_aero*=9)
 - Uses Tegen for aerosols not computed by ART
- CAMS climatology
 - *irad_aero*=7, *cams_aero_filename*
 - Create climatology on your grid: `scripts/preprocessing/make_camsclim_onICONgrid.sh`
- Kinne, stratospheric volcanic, simple plumes. Powerful in combination
 - Kinne: constant or time series data
 - Volcanic stratospheric aerosol for CMIP6
 - Simple plumes: possible emissions
 - Useful for climate application
- CAMS 3D forecasted aerosol
 - *irad_aero* =8, *cams_aero_filename*

CAMS climatology	Tegen Climatology
3D climatology of mixing ratios	2D climatology of AOD
5 species of aerosols used (of 9)	5 species of aerosols
Vertical profile based on dynamics	Fixed vertical profile
Optical properties calculated for each RRTM/ecRad WL intervals	Optical properties calculated at 550 nm and corrections made for other WL
Optical properties of hydrophilic aerosols are RH-dependent	Optical properties are RH-independent
Longwave scattering included	longwave scattering not included

- climatology in external parameter file (“extpar”)
- depends on soil type, month, wavelength range, modified for soil moisture, snow, sea ice

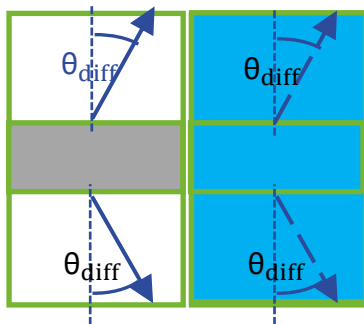
Modelling Radiation

Solver and cloud geometry



Simplifications:

- ignore phase, polarisation
 - only treat up-/downward flux instead of radiances in all directions (2 streams)
 - scattering phase function described by one parameter: asymmetry factor g
 - cloudy and clear region of gridbox
-
- Direct solar radiation is handled separately;
 - Diffuse radiation: assume solar zenith angle θ_{diff} to approximate integral over angles

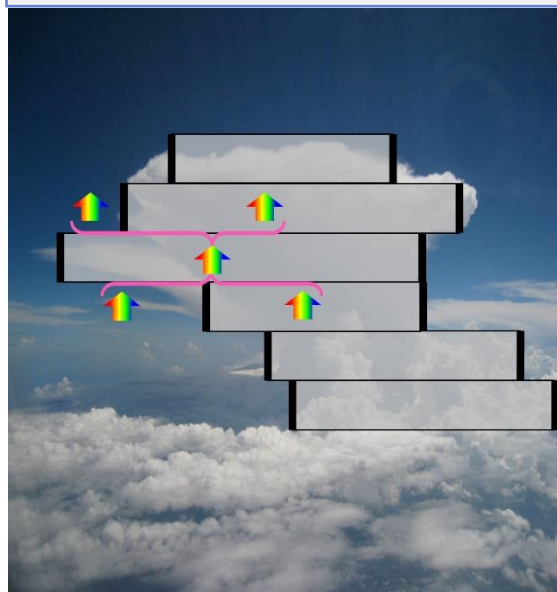


In every layer, for clear and cloudy, compute:

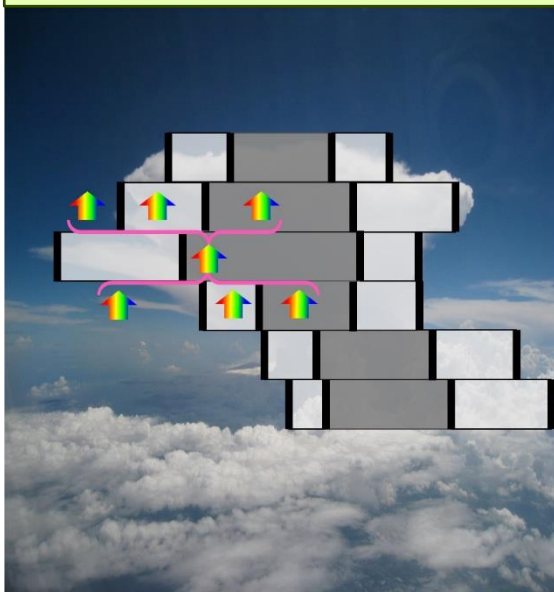
- Extinction (loss)
- Scattering (loss and gain)
- Internal source (gain)

All solvers for global models **simplify** by treating **only vertical** dimension explicitly.

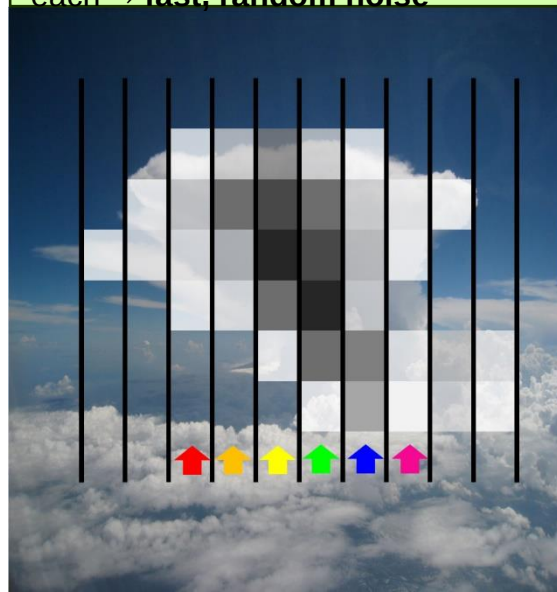
Two-stream solver (e.g. RRTM in ICON): solve in **cloudy / clear regions**, partition at layer boundaries according to **overlap**



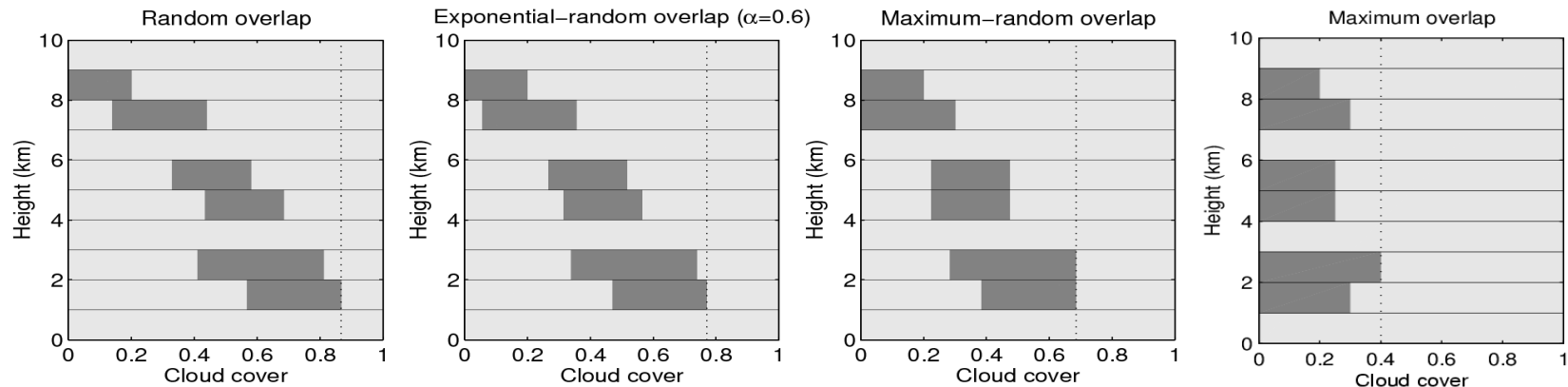
Tripleclouds/SPARTACUS (ecRad): similar; 3 regions: **clear, thin cloud, thick cloud** → subgrid cloud inhomogeneity



McICA (ecRad): draw **random clouds in sub-columns** for overlap + inhomogeneity; **distribute spectral intervals** in 1 sub-column each → **fast, random noise**



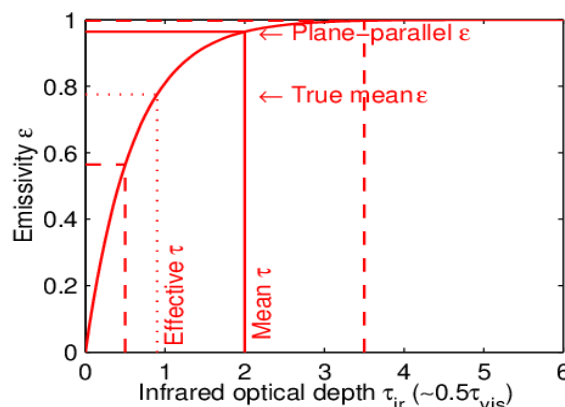
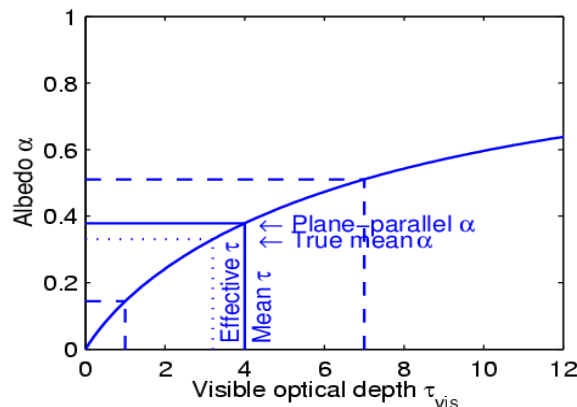
- For given cloud fraction in each layer, **cloud overlap** determines total cloud cover



Adapted from Hogan & Illingworth 2000, QJRMS

- Based on observations (Hogan & Illingworth 2000): **exponential-random overlap**, decorrelation length ca. 2km, or dependent on cloud type
- Development: latitude-dependent decorrelation length-scale

Reflectivity and longwave emissivity are **non-linear** functions of optical depth / cloud water
 → mean optical properties $\neq$ optical properties of mean optical depth τ
 Using reflectivity and emissivity of mean optical depth overestimates cloud effect.



Plot by R. Hogan

- ecRad: sub-divide cloud into two or more sub-regions with different optical depth (Pincus et al. 2003; Shonk & Hogan 2008)
- Assumed distribution (Gamma/log-normal), inhomogeneity parameter $\text{FSD} = \sigma/\text{mean} = 1$

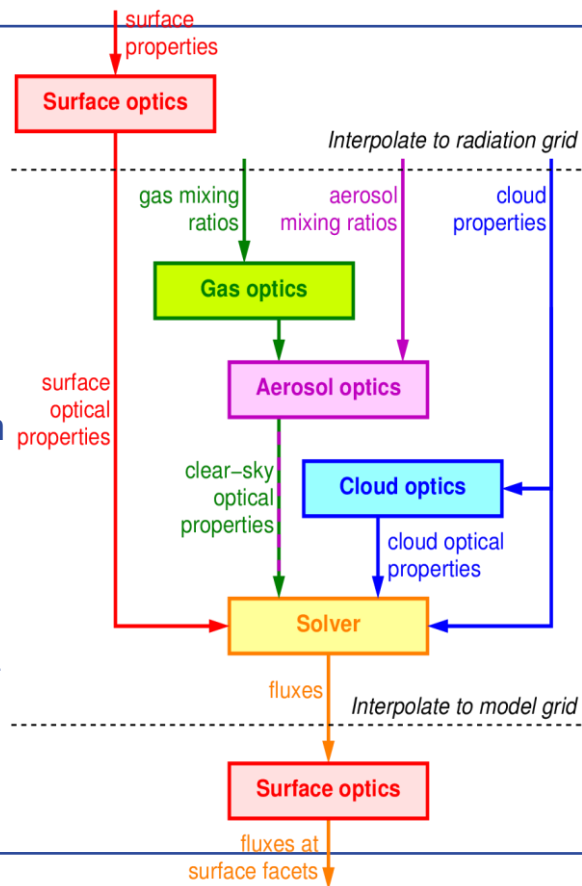
- All solvers simplify by mainly considering up/down flux within the model column
- Greatest uncertainty stems from interaction with cloud!
 - How to partition clear/cloudy areas in grid box
 - How to arrange clear/cloudy areas vertically (overlap)
 - How to represent subgrid-variability more generally
 - Cloud microphysics also uncertain
- Note: RRTM is a solver (old, two-stream), RRTMG is a gas-optics package!

ICON Radiation Scheme

ecRAD, and how to run it

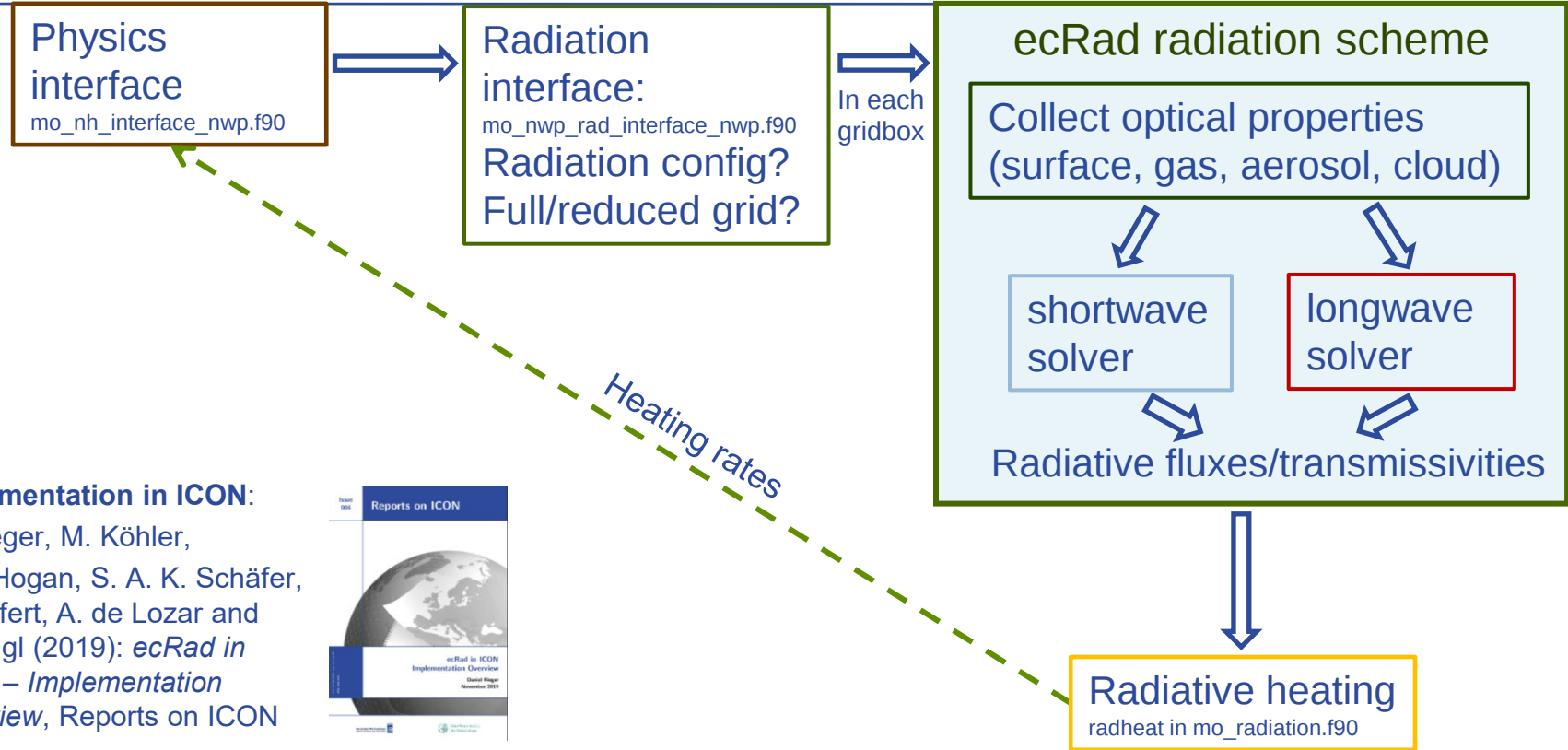


- Aerosol optics: variable species number and properties (set at run-time)
- Gas optics
 - RRTMG (Iacono 2008)
 - ecCKD (Hogan 2010): fewer spectral intervals needed for similar precision
- Cloud optics:
 - liquid: SOCRATES (MetOffice), Slingo (1989)
 - ice: Fu 1996, 1998 (default), Yi et al. 2013 or Baran et al. 2014
- Surface (under development)
Consistent treatment of urban and forest canopies



Solvers for radiative transfer equations:

- McICA (Pincus et al. 2005), Tripleclouds (Shonk & Hogan, 2008) or SPARTACUS (Schäfer et al. 2016, Hogan et al. 2016)
- Longwave scattering optional
- Can configure cloud overlap
- Cloud inhomogeneity: can configure width and shape of PDF

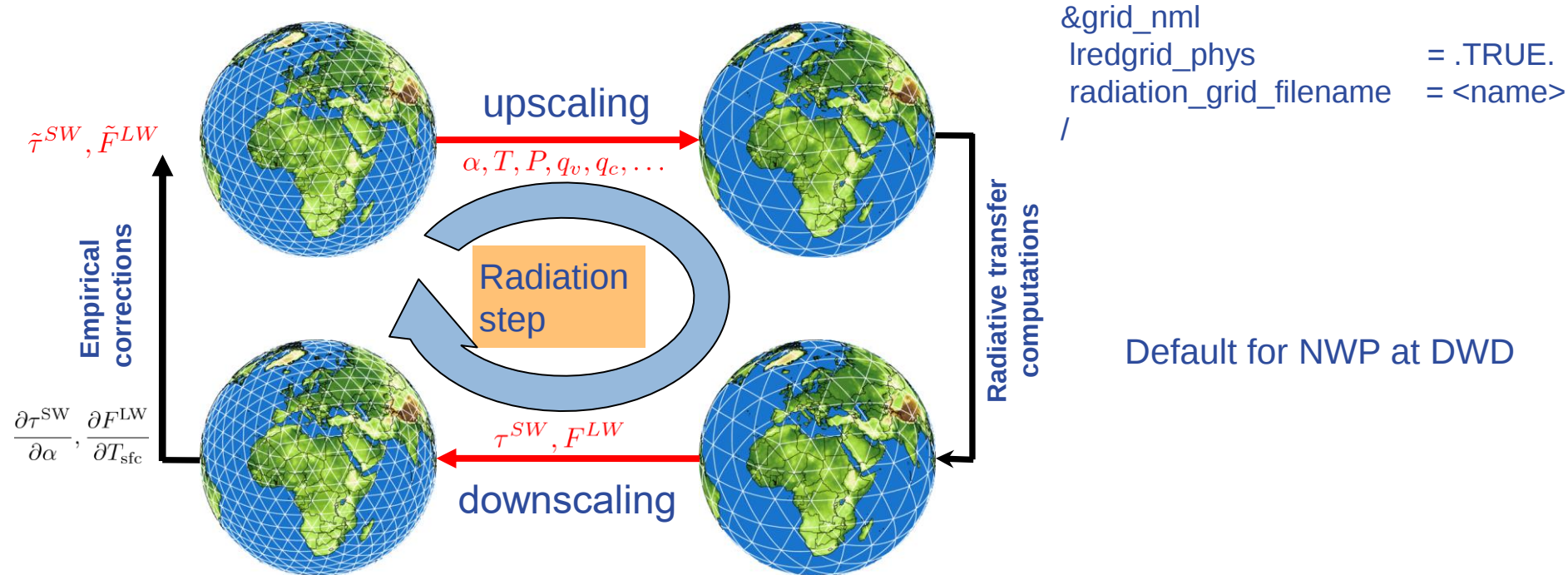


Implementation in ICON:

D. Rieger, M. Köhler,
R. J. Hogan, S. A. K. Schäfer,
A. Seifert, A. de Lozar and
G. Zängl (2019): *ecRad in
ICON – Implementation
Overview*, Reports on ICON



Hierarchical structure of triangular mesh: calculate physical processes (e.g. radiative transfer) with different spatial resolution compared to dynamics.



- Advantages of the reduced radiation grid
 - Fully supported
 - `&nwp_phy_nml: latm_above_top`
→ reduce TOA bias by adding an extra layer
 - Scores better than with full grid
 - Accurate
 - Interpolation quicker than running radiation on full grid
 - Load balancing available: mix day and night time columns on every processor
 - Radiation time step larger than physics time step, so reducing spatial resolution is reasonable

Use the reduced radiation grid!

configure: `./configure --enable-ecrad` (default)

&nwp_phy_nml

`inwp_radiation = 4`

`dt_rad = ?`

! 0: no radiation, 1: RRTM, 4: ecRad

! Radiation time step. Use multiple of convection time step!

&radiation_nml

`ecRad_data_path = <ICON-directory>/externals/ecrad/data`

`icld_overlap = 2`

`irad_aero = 6`

`ecrad_isolver = 0`

`ecrad_igas_model = 0`

`ecrad_liquid_scatter = 0`

`ecrad_ice_scatter = 0`

`ecrad_llw_cloud_scatter = .false.`

! 1: maximum-random, 2: exponential-random, 3: maximum, 4: random

! Aerosols; 0: no aerosol, 2: constant, 6: Tegen climatology, 9: ART

! Radiation solver; 0: McICA, 1: Tripleclouds (Hogan and Shonk 2008)

! Gas model and spectral bands; 0: RRTMG, 1: ecckd (Hogan and Matricardi 2020)

! Liquid optics scheme: 0: SOCRATES, 1: Slingo (1989)

! Ice optics scheme: 0: Fu et al. (1996), 1: Baran et al. (2016)

! Do longwave cloud scattering?

Not all combinations possible. ecRad documentation at <https://confluence.ecmwf.int/display/ECRAD>

Current operational: RRTMG gas optics + McICA solver

Alternative: ecCKD gas optics + TripleClouds solver → faster on other architectures

McICA solver with ecCKD is noisy with only 64 g-points. Not recommended.

DO NOT USE RRTM RADIATION!
Buggy, outdated, no longer supported.

- Radiation drives weather and climate
- Described fully by Maxwell equations + gas emission / absorption
- **Simplification** in global models: **two-stream equations** for up-/downward flux
- Cloudy + clear region in gridbox, bulk optical properties
- Spectrum divided into bands to capture variable emission and absorption
- Some uncertainties, especially in clouds (particle size, geometry,...)
- **ICON:** modular **ecRad:** fast and flexible radiation scheme, parametrisations can be changed individually.
 - Treats cloud inhomogeneity and can treat 3D cloud effects

Thank you for your attention!



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Questions?



■ Radiation textbooks

- Petty G. W., 2006. *A first course in atmospheric radiation*. Sundog Pub.
- Liou, K.-N., 1992: *Radiation and Cloud Processes in the Atmosphere*. Oxford University Press.
- **ecRad**: Hogan, R. J., Bozzo, A. (2018). *A flexible and efficient radiation scheme for the ECMWF model*. JAMES.
- **ecRad in ICON**: Rieger, D., M. Köhler, R. J. Hogan, S. A. K. Schäfer, A. Seifert, A. de Lozar and G. Zängl (2019). *ecRad in ICON - Implementation Overview, Reports on ICON*.
- **RRTM gas optics**: Mlawer, E. J. et al., 1997. *Radiative transfer for inhomogeneous atmospheres: RRTM, a validated correlated-k model for the longwave*. JGR.
- **Cloud geometry**: Shonk, J. K.P. et al., 2010. *Effect of improving representation of horizontal and vertical cloud structure on the Earth's global radiation budget. Part I + Part II*. QJRMS.

